

## **SCIENCE TEAM ORIENTATION & FALL 1999 SEMESTER REVIEW**

- Summary of Science Mission for New & Prospective Team Members (6 minutes)
- Review of Previous Progress (4 minutes)
- Fall 1998 Lightning Activities (6 minutes)
- Fall 1998 Sprite Activities (3 minutes)
- Fall 1998 Photometry Activities (7 minutes)
- Fall 1998 Instrument Design Activities (3 minutes)

## REVIEW OF PREVIOUS PROGRESS

- Objective: build and orbit a satellite that performs meaningful science
- “Zeroth order” baseline design presented on September 3, 1997 and intended to:
  - 1) provide the other teams with a starting point for their work
  - 2) provide subsystem teams with approximate, size, mass & volume constraints
  - 3) provide a “benchmark” for evaluation of alternative mission concepts
- Prototype sensor designed and built to:
  - 1) provide subsystem teams with design details not available from baseline design
  - 2) test certain of the baseline design flight hardware algorithms
  - 3) provide students with requested “hands-on” experience
- Semester Review in May, 1998, provided subsystem teams with:
  - 1) Approximate Flux Values
  - 2) Approximate Data Rates
  - 3) Preliminary Data Format
  - 4) Preliminary Bus Voltages
  - 5) Clarified Science Objectives

## UASAT “ZEROTH ORDER” BASELINE DESIGN VS. OTHER LIGHTNING SENSORS

|                            | LIS                            | OTD                            | UASAT Baseline               |
|----------------------------|--------------------------------|--------------------------------|------------------------------|
| <b>Optics</b>              | <b>Wide-Field-of-View Lens</b> | <b>Wide-Field-of-View Lens</b> | <b>Reflector</b>             |
| <b>Field of View</b>       | <b>80x80 degrees</b>           | <b>100x100 degrees</b>         | <b>5x5 degrees</b>           |
| <b>Field of View</b>       | <b>600x600 km</b>              | <b>1300x1300 km</b>            | <b>1100x200 km</b>           |
| <b>Sensor Size</b>         | <b>8 inch diam x 14 inch</b>   | <b>8 inch diam x 15 inch</b>   | <b>6 inch diam x 20 inch</b> |
| <b>Spatial Resolution</b>  | <b>5-10 km</b>                 | <b>10 km</b>                   | <b>10 km</b>                 |
| <b>Wavelength</b>          | <b>777.4 nm</b>                | <b>777.4 nm</b>                | <b>All Visible</b>           |
| <b>Detector</b>            | <b>Photodiode Array</b>        | <b>Photodiode Array</b>        | <b>CCD</b>                   |
| <b>Detector Size</b>       | <b>128 x 128</b>               | <b>128 x 128</b>               | <b>TBD</b>                   |
| <b>Temporal Resolution</b> | <b>2 ms</b>                    | <b>2 ms</b>                    | <b>TBD</b>                   |
| <b>Power</b>               | <b>25 watts</b>                | <b>70 watts</b>                | <b>TBD</b>                   |

## FALL 1998 LIGHTNING ACTIVITIES

- Began Detailed Investigations of OTD & LIS.
- Began Investigation of Polar Lightning Observations
- Developed Mission Operation Timeline & Protocols
- Continued Development of Detailed Lightning Detection Algorithm
- Located Experimental Prototype CCD Optimized for Lightning Detection
- Began Trade-off Studies and Definition of Alternative Mission Concepts
  - Statistical Significance of Lightning Counts vs. Sensor Field of View
  - Readout Speed vs. Statistical Significance of Lightning Counts
  - Power Consumption vs. Readout Speed
  - Power Consumption vs. Observation Duration
  - Observation Duration vs. Statistical Significance of Lightning Counts
  - Readout Speed vs. Processor Size
  - Processor Size vs. On-Board Storage of All Images
  - Processor Size vs. Real-time Downlink
  - CCD vs. Photodiode Array
  - Detector Dynamic Range vs. Statistical Significance of Lightning Counts

## FALL 1998 SPRITE ACTIVITIES

- Field Tested Prototype in Yucca Ridge Colorado
  - Identified Necessary Improvements: Wide FOV Guiding Camera
  - Mounting and Pointing Improvements
  - CCD Exposure Trigger
- Designed Laboratory Test Apparatus for “Simulated Sprites”
- Continued Development of Detailed Sprite Detection Algorithm
- Began Investigation of Statistical Sampling Requirements for Scientific Merit
- Began Trade-off Studies:
  - Detector Dynamic Range vs. Statistical Significance of Spite Counts
  - Scientific Merit vs. Image Resolution
  - CCD vs. Photodide Array
- Began Investigation of Alternative Sprite Detection Concepts:
  - Separate Intensified CCD Camera
  - Electronic vs. Optical Image Processing (Power Consumption)
  - RF Signature Detection
- Began Investigation of Alternative Air Glow/Aurora Discrimination Concepts:

## FALL 1998 PHOTOMETRY ACTIVITIES

- Completed Design and Construction of Basic Optical and Electronic Prototype  
-- David Sing & Guthrie Partridge
- Began Design of Filter Wheel and Aperture Wheel
- Interfaced Prototype Detector to Data Acquisition Board via LabView Software
- Began Definition of Target List and Filter Selection
- Began Trade-off Studies:
  - Exposure Times & Detector Sensitivity vs. Scientific Merit
  - Scientific Merit vs. Aperture Size
  - Scientific Merit vs. Pointing Accuracy and Stability
  - Power Consumption & Complexity vs. Internal Pointing Correction (Fold-Mirror)
- Began Studies to Quantify Accuracy Goals and Calibration Methodology
- Began Feasibility Study for Extra-Solar Planet Detection

## FALL 1998 INSTRUMENT DESIGN ACTIVITIES

- Obtained Machine Shop Training for 6 Team Members
- Prepared Recruiting Letters In August, 1998 to Meet Need for Optical and Electrical Engineering Students
- Began Trade-off Studies:
  - Optimization for Photometry vs. Lightning Detection
  - Optical Speed vs. Field of View
  - Reflector vs. Refractor
  - Resolution of Most of These Issues Dependent on Results of Lightning, Sprite and Photometry Studies
- Began Investigation of Wynn Inverse Cassegrain Design (7 degree FOV Reflector)
- Struggled with the task of developing a single sensor for three independent science experiments